



# Development status of the Second-generation Global Imager (SGLI) on GCOM-C1



**September 2, 2009**  
**SPIE Europe, Berlin**



**NEC/TOSHIBA**  
NEC東芝スペースシステム

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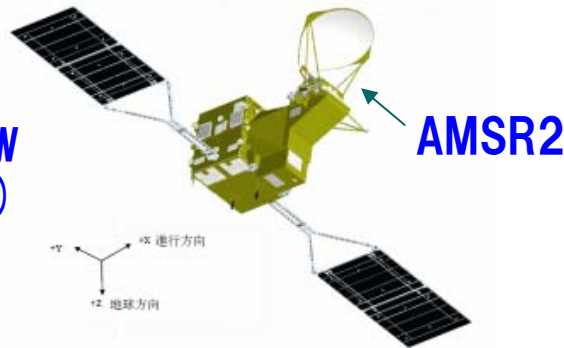
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- SGLI on GCOM-C1 Satellite
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# Global Change Observation Mission (GCOM)

- Global observation satellite system as JAXA's GEOSS contribution.
- 2 satellite series for 5 years, total 13 years observation.
  - ✓ GCOM-W Microwave radiometer observation for WATER CYCLE using AMSR2 (AMSR-E follow on)
  - ✓ GCOM-C Optical multi-channel observation for RADIATION BUDGET and CARBON CYCLE using SGLI (GLI follow on)

**GCOM-W  
(WATER)**

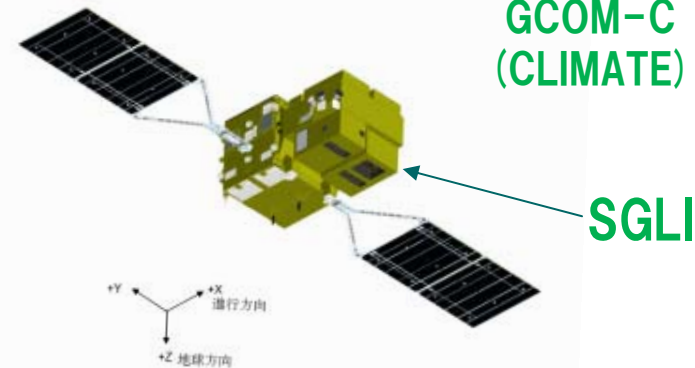


**Advanced Microwave Radiometer 2  
(AMSR2)**

Sensor

Passive Microwave Observation  
 Water vapor, soil moisture etc

**GCOM-C  
(CLIMATE)**

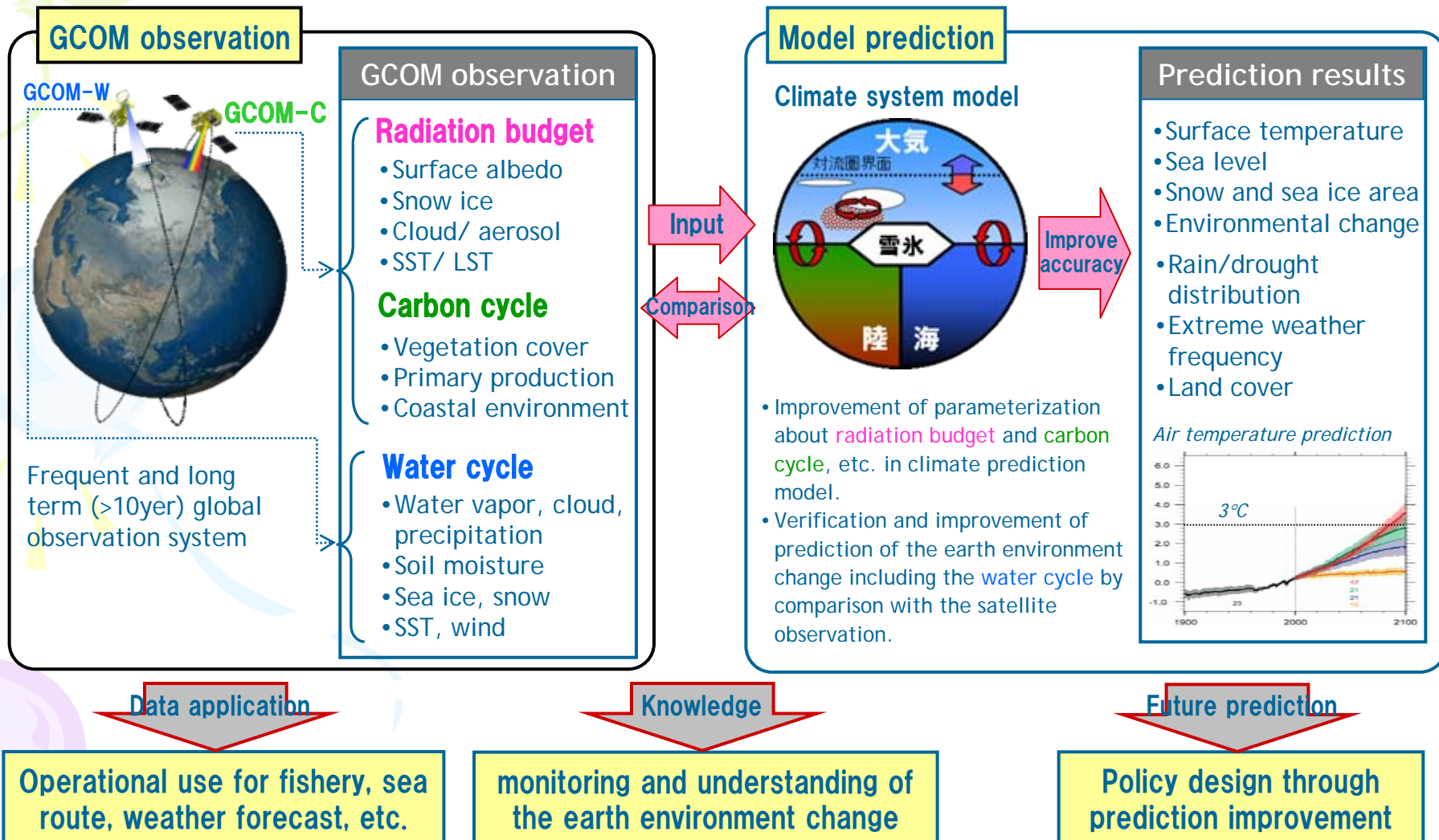


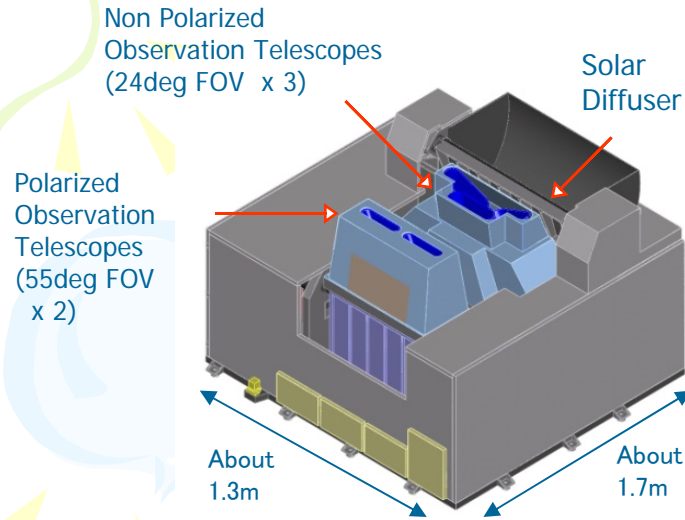
**Second Generation Global Imager  
(SGLI)**

Sensor

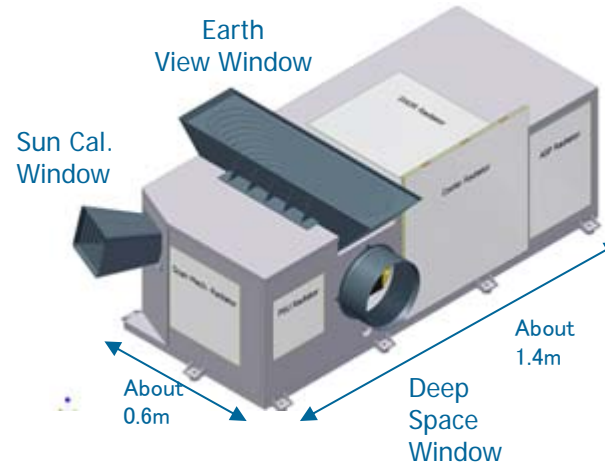
Optical Observation 380nm – 12 micron  
 Cloud, Aerosol, Vegetation, Chlorophyll etc

# GCOM Mission

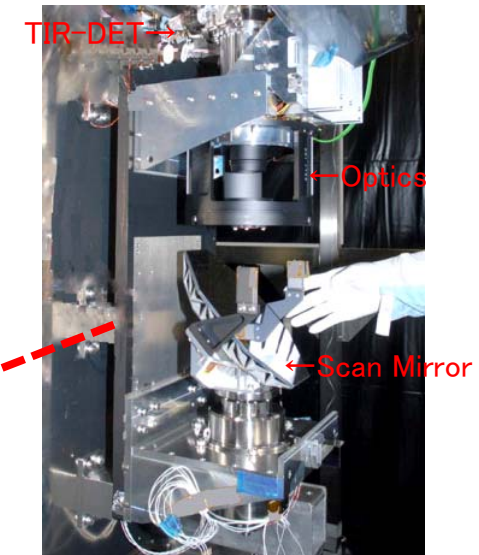




Visible and Near Infrared Radiometer (SGLI-VNR)



Infrared Scanning Radiometer (SGLI-IRS)



SGLI IRS Bread Board Model

Sensor Unit	features
SGLI VNR	Non Polarized Observation (11ch), IFOV 250m, Swath 1150km Polarized Observation (2ch), IFOV 1km, Swath 1150km
SGLI IRS	Shortwave Infrared (SWI 4ch), IFOV 250m/1km, Swath 1400km Thermal Infrared (TIR:2ch), IFOV 500m, Swath 1400km

# SGLI Specification

- The SGLI features are 250m (VNI) and 500m (T) spatial resolution and polarization/along-track slant view channels (P), which will improve land, coastal, and aerosol observations.

*250m over the Land or coastal area, and 1km over offshore*

GCOM-C SGLI characteristics	
Orbit	Sun-synchronous (descending local time: 10:30) Altitude 798km, Inclination 98.6deg
Mission Life	5 years (3 satellites; total 13 years)
Scan	Push-broom electric scan (VNR) Wisk-broom mechanical scan (IRS)
Scan width	1150km cross track (VNR: VN & P) 1400km cross track (IRS: SW & T)
Digitalization	12bit
Polarization	3 polarization angles for P
Along track direction	Nadir for VN, SW and T, +45 deg and -45 deg for P
On-board calibration	VN: Solar diffuser, LED, Lunar cal maneuvers, and dark current by masked pixels and nighttime obs. SW: Solar diffuser, LED, Lunar, and dark current by deep space window T: Black body and dark current by deep space window

*Multi-angle obs. for 674nm and 869nm*

SGLI channels						
CH	$\lambda$	$\Delta\lambda$	$L_{std}$	$L_{max}$	SNR at Lstd	IFOV
	VN, P, SW: nm T: $\mu\text{m}$		VN, P: W/m <sup>2</sup> /sr/ $\mu\text{m}$ T: Kelvin		VN, P, SW: SNR T: NE $\Delta$ T	m
VN1	380	10	60	210	250	250
VN2	412	10	75	250	400	250
VN3	443	10	64	400	300	250
VN4	490	10	53	120	400	250
VN5	530	20	41	350	250	250
VN6	565	20	33	90	400	250
VN7	673.5	20	23	62	400	250
VN8	673.5	20	25	210	250	250
VN9	763	8	40	350	1200	250/1000
VN10	868.5	20	8	30	400	250
VN11	868.5	20	30	300	200	250
P1	673.5	20	25	250	250	1000
P2	868.5	20	30	300	250	1000
SW1	1050	20	57	248	500	1000
SW2	1380	20	8	103	150	1000
SW3	1630	200	3	50	57	250
SW4	2210	50	1.9	20	211	1000
T1	10.8	0.7	300	340	0.2	250/500
T2	12.0	0.7	300	340	0.2	250/500

*option*

# Bread Board Model (BBM) Activity

## Visible and Near infrared radiometer (VNR)

**Critical Technologies**

- To Validate feasibility of the critical components



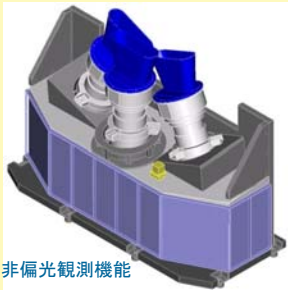
• Bandpass filters



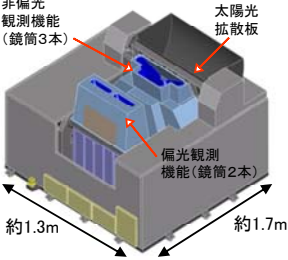
• CCD

**BBM System Design**

- To establish SGLI system design to achieve observation requirement



非偏光観測機能






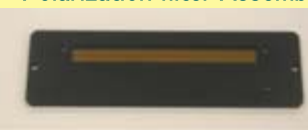


太陽光拡散板  
 偏光観測機能(鏡筒2本)  
 非偏光観測機能(鏡筒3本)  
 約1.3m      約1.7m

**BBM Component Tests**

- To assess and demonstrate the observation performances by component level

<Major BBM component>

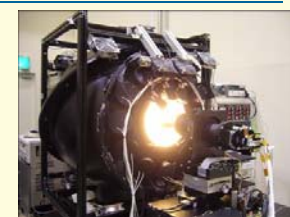
- 11line CCD
- Bandpass filter Assembly
- Optics
- Polarization filter Assembly
- Readout Circuit

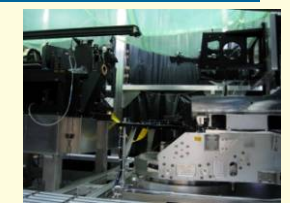
**BBM System Tests**

- To assess and demonstrate the observation performances by system level


**Radiometric Tests**



**Geometric Tests**



**Tests with onboard calibrators**



**Complete**

**Complete**

**Complete**

**Complete**

# Bread Board Model (BBM) Activity Infrared Scanning radiometer (IRS)

**Critical Technologies**

- To Validate feasibility of the critical components



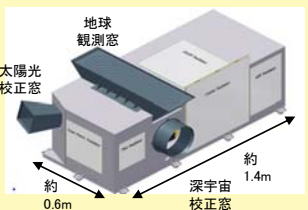
• Mirror Sample



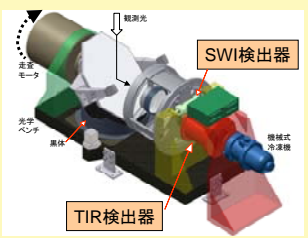
• TIR detector

**BBM System Design**

- To establish SGLI system design to achieve observation requirement



地球観測窓  
 太陽光校正窓  
 約 0.6m  
 約 1.4m  
 深宇宙校正窓







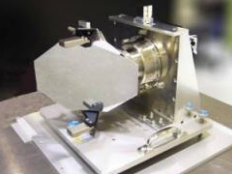


観測光  
 SWI検出器  
 TIR検出器  
 機械式冷却機  
 光学ベンチ  
 異体  
 基板モジュール

**BBM Component Tests**

- To assess and demonstrate the observation performances by component level

**<Major BBM component>**


- SWIR Detector
- TIR Detector
- Optics
- Mechanical Cooler
- Scan Mirror Assembly
- Readout Circuit
- Blackbody


**BBM System Tests**

- To assess and demonstrate the observation performances by system level

**SWIR Radiometric Tests**



**TIR Radiometric Tests (Vacuum condition)**



**Geometric Tests**

**Tests with onboard calibrators**

**Complete**

**Complete**

**Complete**

**Complete**

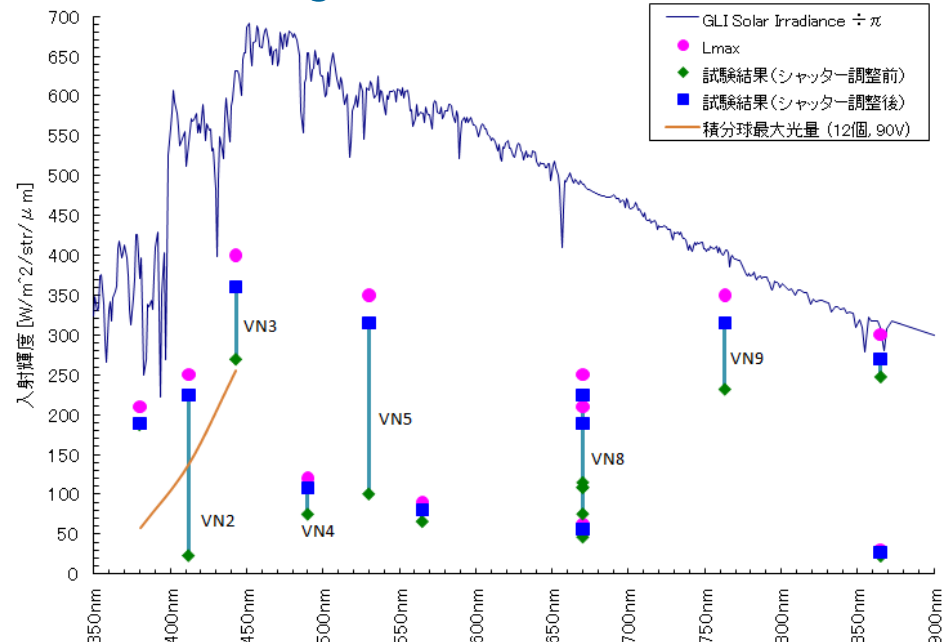
Development status of SGLI at SPIE Europe, Berlin 2009



# BBM test results of VNR

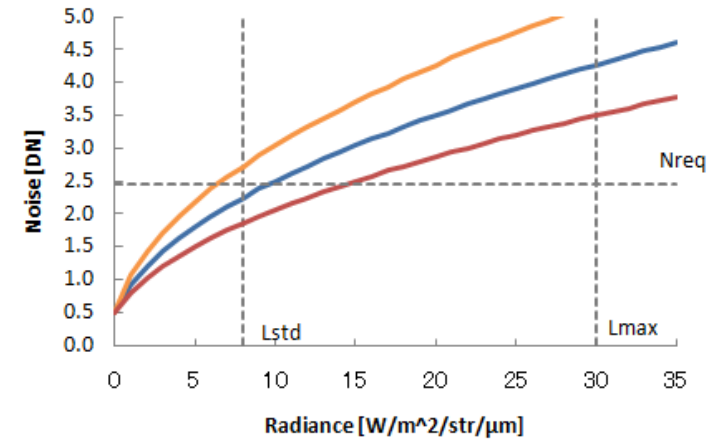
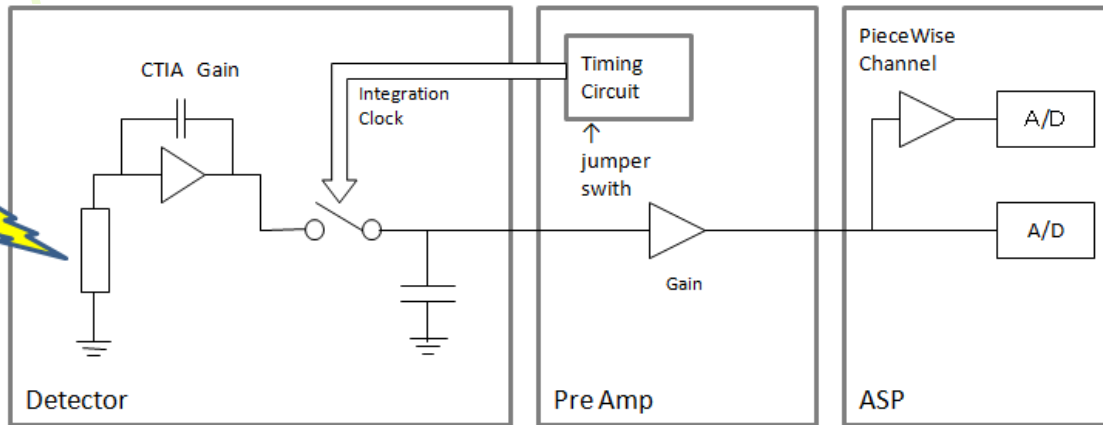
## Dynamic Range

- VNR dynamic range requirement is set based on careful study of target characteristics.
- VNR uses CCD. Integration time can be easily changed by command.
- This means dynamic range adjustment is possible.
  - ✓ Difference between design and manufacturing results such as transmittance of optics can be adjusted
  - ✓ Dynamic range modification can be changed even after launch
  - ✓ Calibration technique using different integration time.

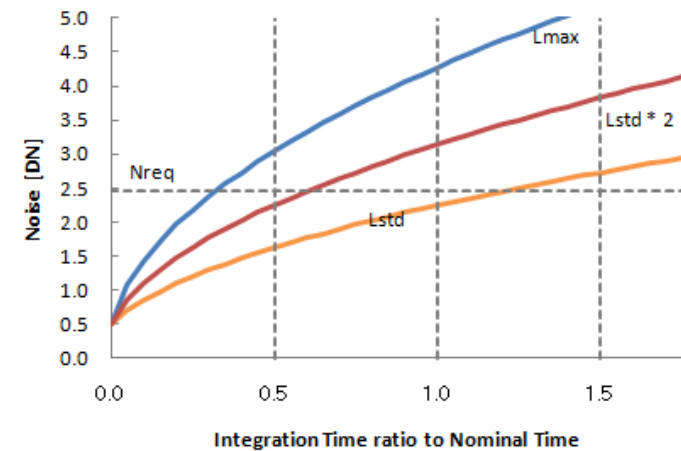
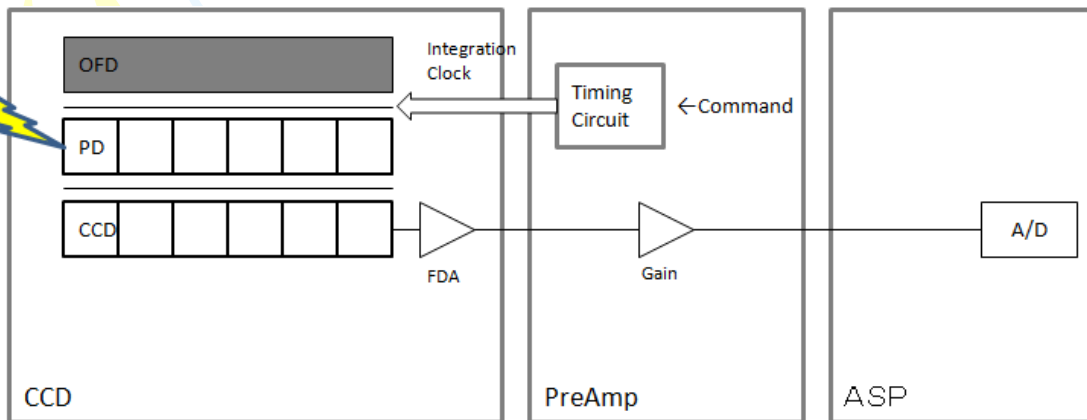


# BBM test results of VNR CCD Integration Time

## Previous GLI design



## SGLI VNR design

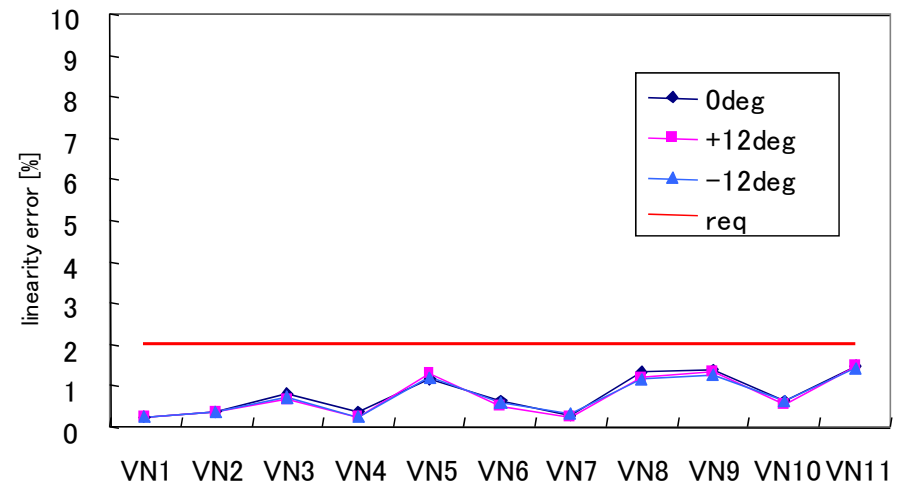
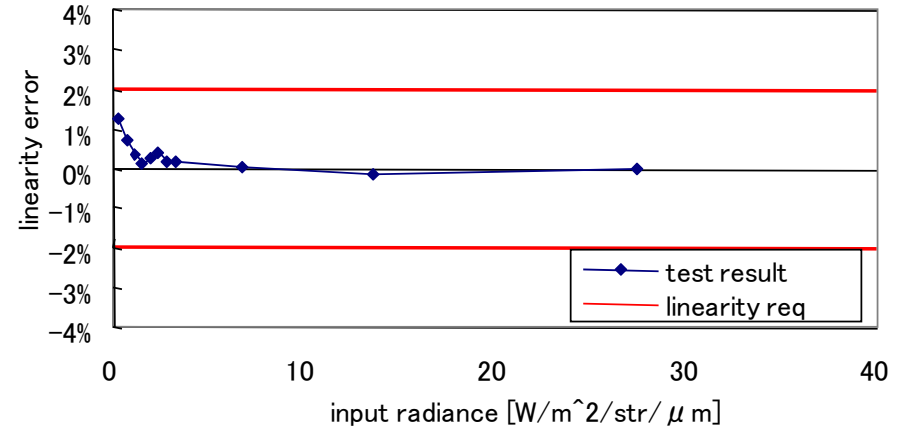


# BBM test results of VNR

## Linearity

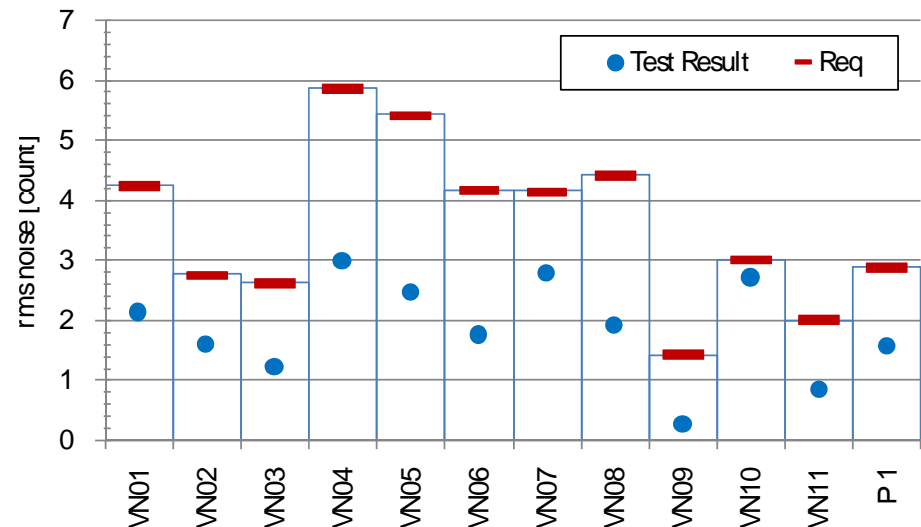
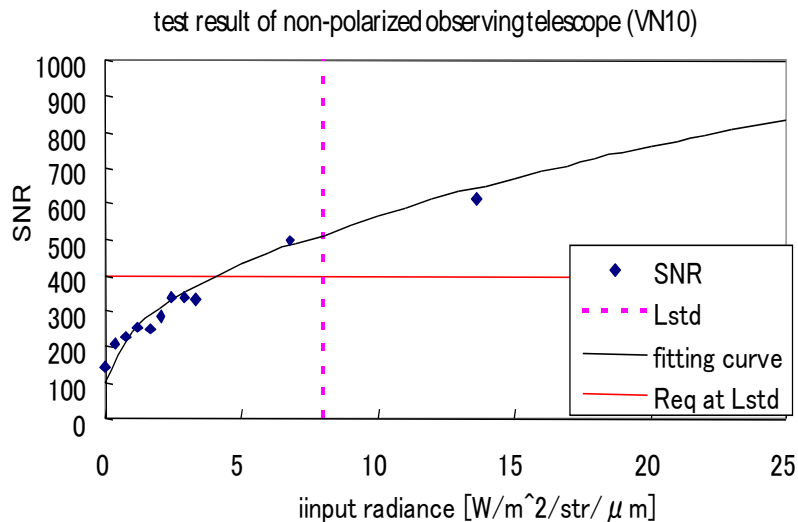
- Linearity is very important for radiometer because it leads to accuracy dependencies to input radiance level.
- Also, the test methodology using integration time above mentioned uses the assumption that sensor is very linear.
- VNR BBM linearity is checked for both input radiance dependencies and integration time dependencies.
- Test result is very good and further detailed analysis is conducted to find the major non-linear element to improve.

test result of VN10 non-polarized observing telescope



# BBM test results of VNR Noise

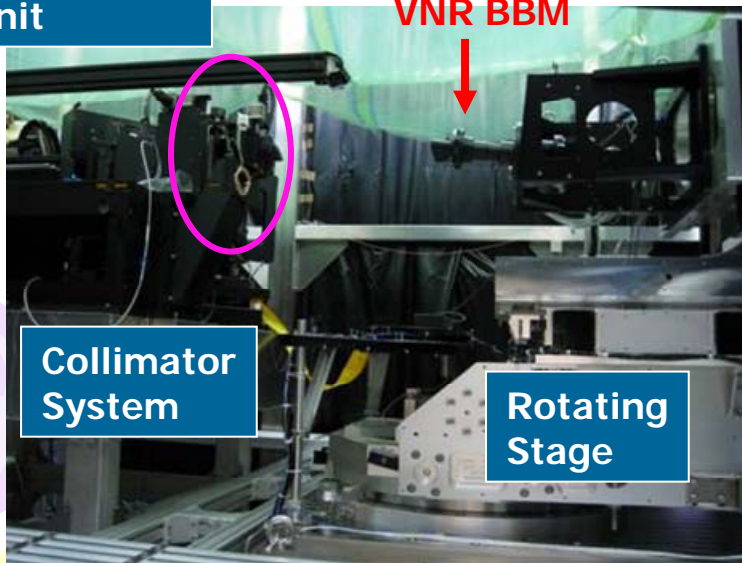
- SGLI noise performance is required as Signal to Noise Ratio (SNR) at standard signal observation radiance (Lstd) for each observation channel.
- Both integration time dependence and input radiance for SNR was checked.
- The all channels noise level met the sensor specification.



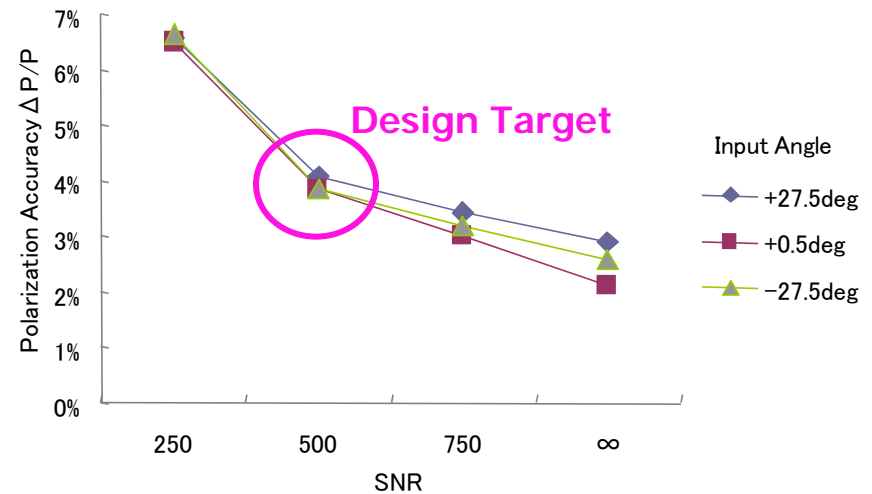
# BBM test results of VNR Polarized Light Observation

- The special polarization unit is set to the collimator light source. The collimator polarization characteristic is canceled with de-polarizer filter and polarized unit can be rotated around the optical axis to realize the desired polarized direction for both linear and partially polarized light.
- As the result of BBM test, approximately 4% polarization observation accuracy ( $\Delta P/P_{std}$ ) and 1.3deg polarization direction error is confirmed for the 670nm polarization observing telescope with standard 5% polarized light ( $P_{std}$ ).

Polarization Unit

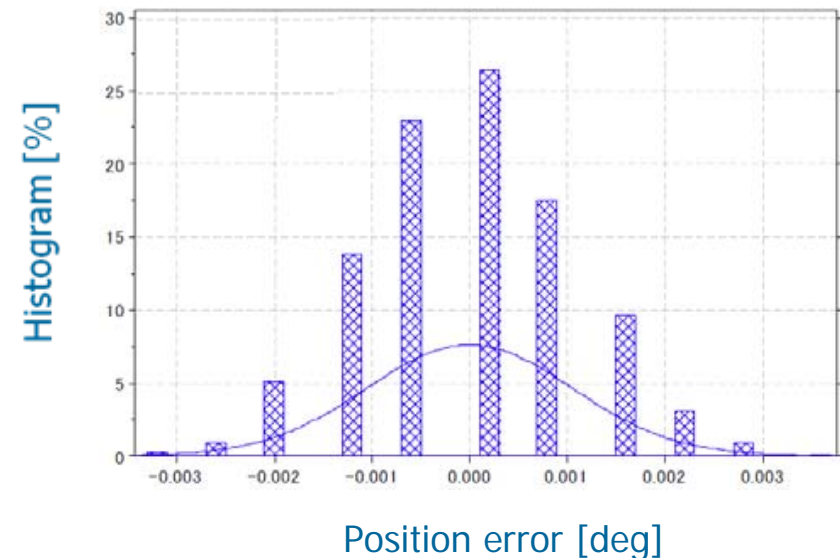
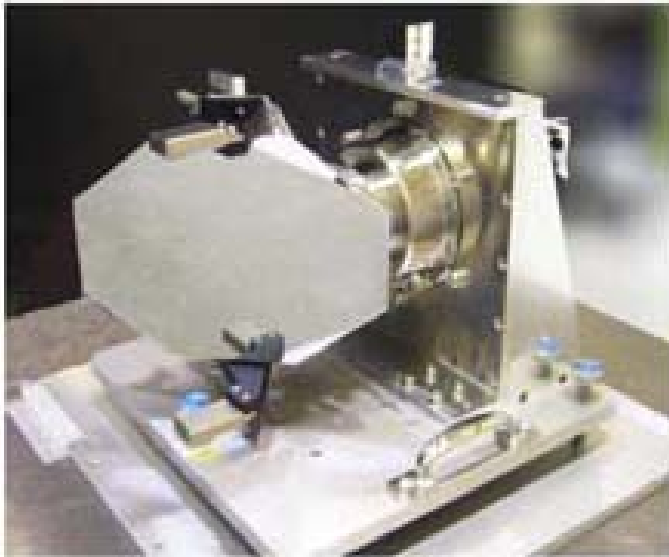


test result of 670P polarization observing telescope



# BBM test results of IRS Scanning Performance

- Because very stable rotating speed is important for sensor observation geometry, the centroid position error and product of inertia should be minimized.
- Manufacturing error and assembly process were carefully studied and confirmed. The friction learning technique in the rotation speed control loop is also applied.
- The selection of the bearing unit is another key issue for the IRS 5years operation requirement. The bearing lifetime test under the vacuum condition was successfully completed for required  $2.7 \times 10^8$  revolution operation and still under the test for the further design margin.



# BBM test results of IRS

## SWI Radiometric Performance

### ■ SNR performance

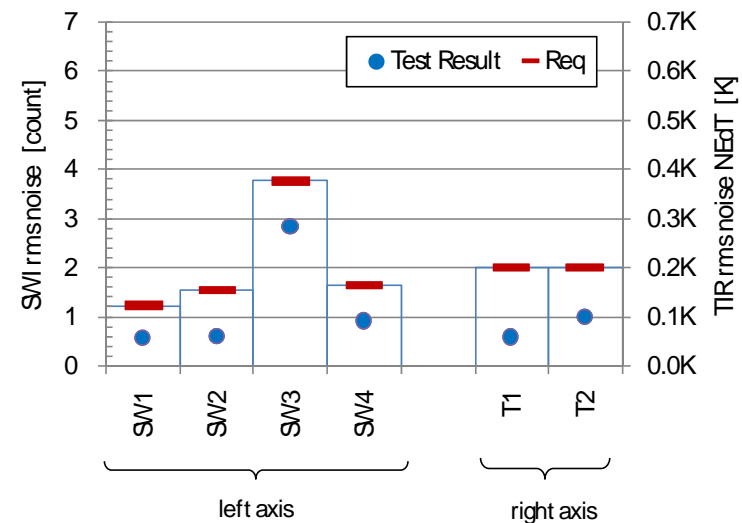
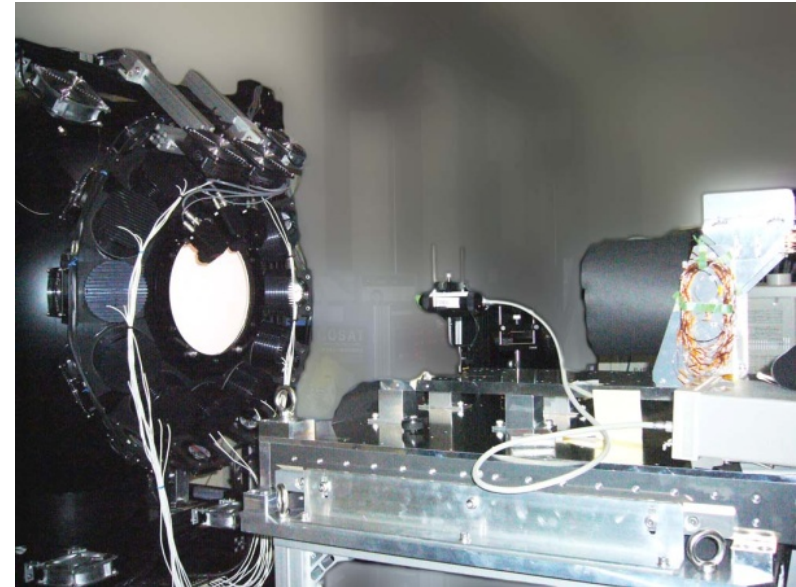
- ✓ high transmittance of telescope and filters
- ✓ detector sensitivity
- ✓ dark current of 2.2  $\mu$  m channel at operating -30deg C
- ✓ System SNR performance

### ■ Dynamic Range

- ✓ adjustment at analog signal processor (ASP)

### ■ EMC allowance limit

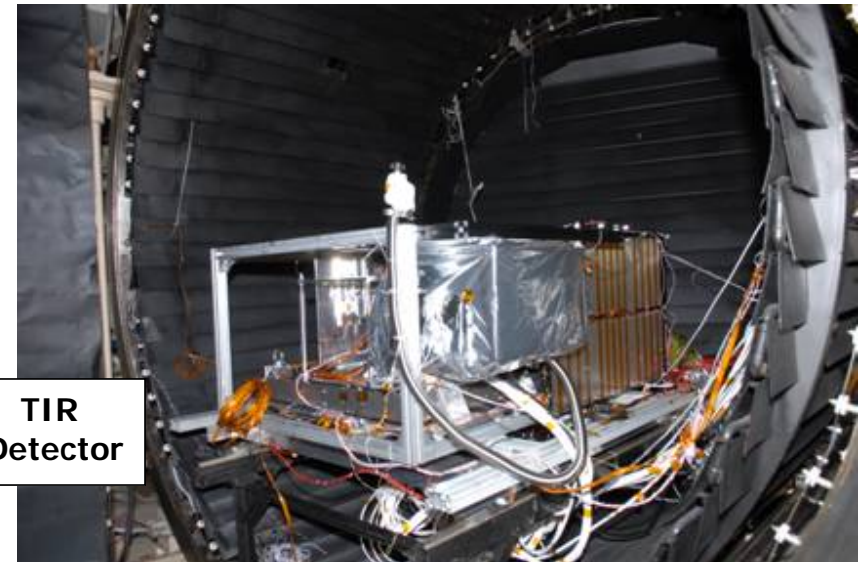
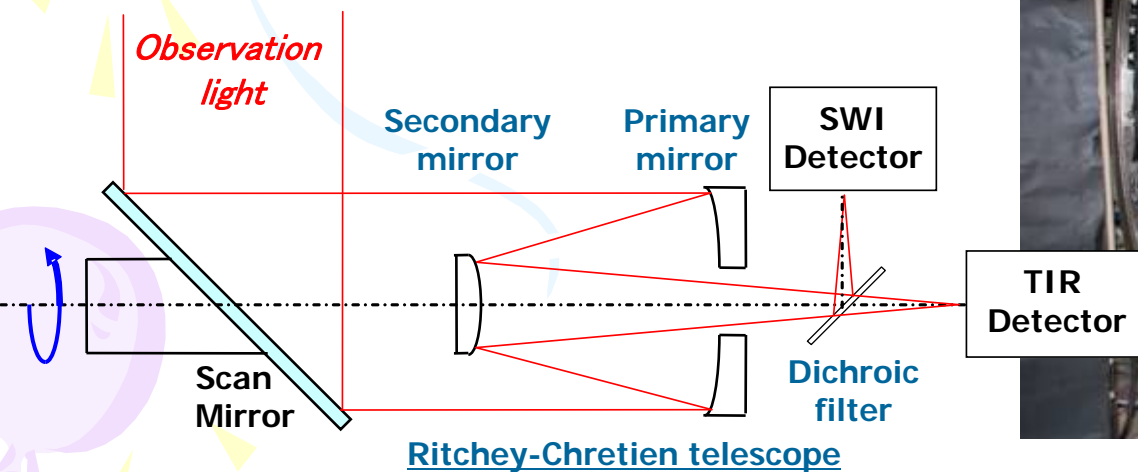
- ✓ Trans Impedance Amplifier (TIA) for detector readout.
- ✓ EMC test in the shield room for both RF radiation and conductive noise allowance



# BBM test results of IRS TIR Radiometric Performance

- Detector temperature & performance
  - special dewar for 55K
  - spectral characteristics at 55K
  - detector linearity
- Effects of sensor temperature around optical path
  - dynamic range design considering sensor temperature range
  - Thermal vacuum test to confirm the design.

- Dynamic range adjustment methodology
  - integrating capacity at ROIC
  - integration time
  - post amplifier gain (fixed)
- TIR system level test meets the noise requirement (NEdT). (previous page)





# SGLI Current Status

- Based on the BBM successful test result the sensor Engineering Model (EM) design is on going.
- The manufacturing of EM starts soon. EM test will be conducted in next year, and sensor PFM manufacturing starts.
- The target launch of GCOM-C1 is JAN 2014.

Japanese Fiscal Year	FY2006	FY2007	FY2008	FY2009	FY2010	FY2011	FY2012	FY2013
Milestone		▲ GCOM-W1 Project start				GCOM-W1 Launch ▲		GCOM-C1 Launch (TBD) ▲
GCOM-W1	▲ SDR	▲ PDR		▲ CDR				
	Phase-A	Phase-B	Phase-C	Phase-D				
GCOM-C1		▲ SDR			▲ PDR	▲ CDR		
	Pre-Phase-A	Phase-A	Phase-B	Phase-C	Phase-D			

# Conclusion

- GCOM is JAXA's GEOSS contribution of global observation for water cycle, radiation budget and carbon cycle.
- SGLI is the optical sensor on GCOM-C1 satellite.
- Bread Board Model (BBM) design and test activity has successfully completed last year.
- The Engineering Model (EM) design is currently conducted.
- The EM testing including environmental test and I/F test with satellite system is planned in next year.