

# ATBD of GCOM-C chlorophyll-a concentration algorithm

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JAXA, EORC, Oct. 2018

## 1. Background and objectives

Chlorophyll-a concentration (Chl-*a*) algorithm has more than 20-year history, and blue-green band ratio algorithm (OCx) which calculate Chl-*a* by 4th order empirical equation (Eq. (1)) is currently used as the standard (O'Reilly et al., 2000).

$$\log_{10}(\text{Chl}a) = c_0 + c_1 r + c_2 r^2 + c_3 r^3 + c_4 r^4, \quad \dots (1)$$

where,

$$r = \log_{10}(\text{MAX}[R_{rs}(443\text{nm}, 490\text{nm}, 530\text{nm})] / R_{rs}(555\text{nm})). \quad \dots (2)$$

Recently, NASA/OBPG cobindly uses Color Index based Algorithm (CIA; Hu et al., 2012) in the low Chl-*a* ranges (Eq. (3)) which can reduce noise caused from satellite sensors.

$$\text{Chl-}a = 10^{(-0.4909 + 191.6590 \times \text{CI})} \text{ (when } \text{CI} \leq -0.0005 \text{ sr}^{-1}), \quad \dots (3)$$

where,

$$\text{CI} = R_{rs}(555\text{nm}) - [R_{rs}(443\text{nm}) + (555-443) / (670 - 443) \times (R_{rs}(670\text{nm}) - R_{rs}(443\text{nm}))]. \quad \dots (4)$$

GCOM-C algorithm is developed based on the empirical algorithms for convenient use with the other sensor products. We derived the OCx coefficients optimized for the SGLI spectral bands and characterized estimation error of the SGLI Chl-*a* product.

## 2. Data

### 2.1. NOMAD data

NASA bio-Optical Marine Algorithm Data set (NOMAD) (Werdell and Bailey, 2005) have been used for OCx coefficient of NASA standard Chl-*a* products (OCTS, SeaWiFS, MODIS, VIIRS, ..). NOMAD includes above-water downward-irradiance,  $E_s(\lambda)$ , upward radiance,  $L_w(\lambda)$ , etc., at wavelengths  $\lambda=443, 455, 465, 489, 510, 520, 530, 550, 555, 560, 565, 570, 590, 619, 625, 665, 670\text{nm}$  with 10-nm band width. Remote-sensing reflectance,  $R_{rs}$ , and normalised water-leaving radiance,  $L_{wn}$ , can be calculated from the  $E_s(\lambda)$  and  $L_w(\lambda)$  by the following equation (5).

$$R_{rs}(\lambda) = L_w(\lambda) / E_s(\lambda) \times \text{foQ}(\theta=0, \theta_0=0, \Delta\phi, \text{Chl}a, \lambda) / \text{foQ}(\theta, \theta_0, \Delta\phi, \text{Chl}a, \lambda) \quad \dots (5)$$

$$L_{wn}(\lambda) = R_{rs}(\lambda) \times F_0(\lambda) \dots (6)$$

where, foQ is bi-directional reflectance factor (the look up tables of Morel and Maritorena, 2001 is used),  $F_0$ , solar irradiance,  $\theta$ ,  $\theta_0$ ,  $\Delta\phi$  are satellite zenith, solar zenith and relative azimuth angles respectively.

### 2.2. In-situ data around Japan

Around Japan, several universities (Hokkaido, Nagoya, Tokyo, Yamanashi, and Tokai Univs.) and institutes (Japan Fisheries Research and Education Agency (FRA) including National Research Institute of Fisheries Science, Seikai National Fisheries Research Institute, Tohoku National Fisheries Research Institute, Hokkaido National Fisheries Research Institute), have provided in-situ data for the GCOM-C algorithm development and validation. GCOM-C standard OC4 algorithm is using only NOMAD data in the pre-launch phase considering compatibility with other sensor products. However their data is used for the algorithm evaluation and will be used for algorithm improvements after the GCOM-C launch.

### 3. IOP model for SGLI-bands

We made a simple inherent optical property (IOP) model based on Gordon et al., 1988, Lee et al., 2002 adjusted to NOMAD IOP and  $R_{rs}$  data, and interpolated NOMAD  $R_{rs}$  to 1-nm spectral resolution using the model.

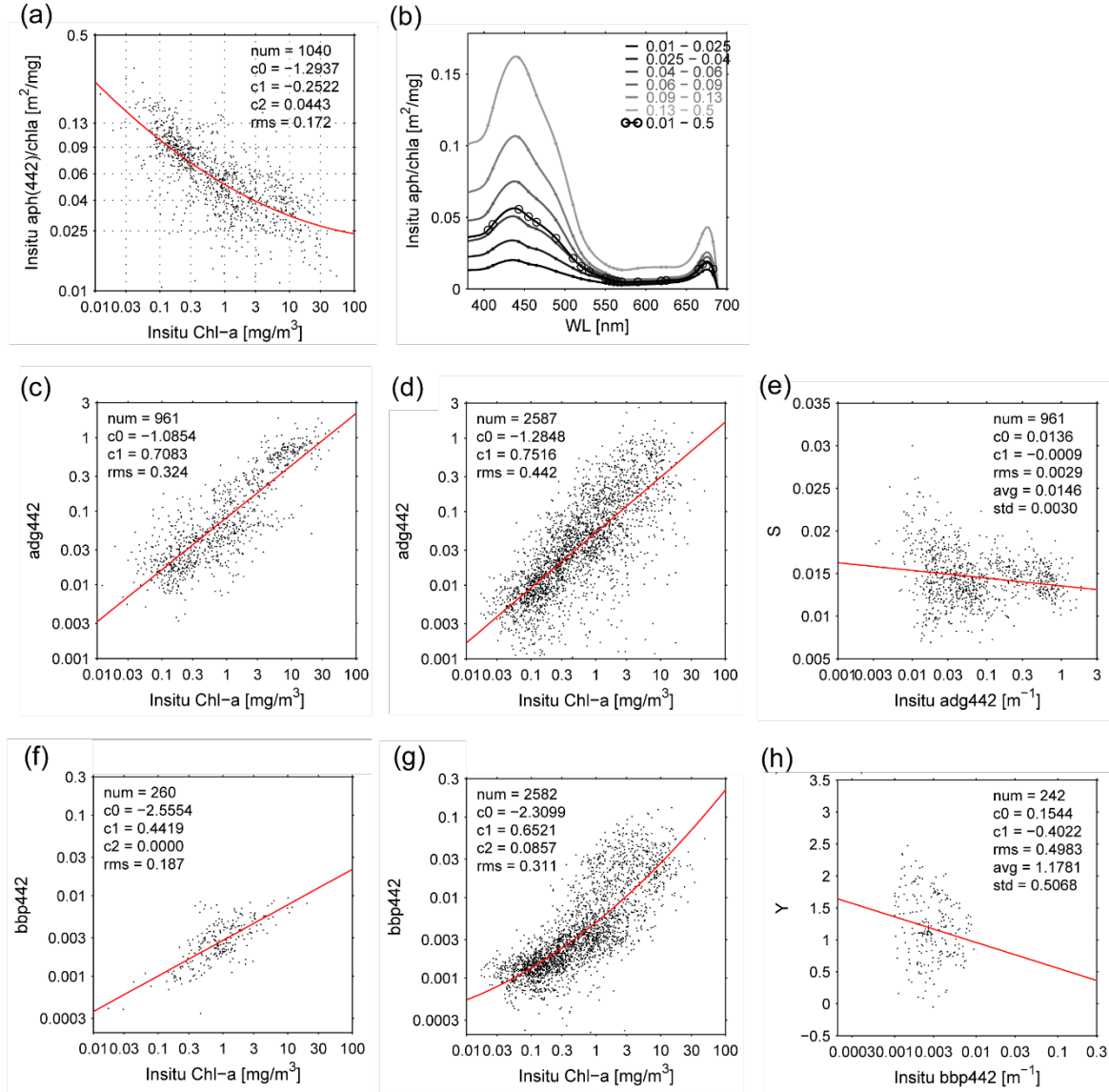


Figure 1. The upper panels are (a) scatter diagram of  $a_{ph}$  and Chl- $a$  and (b) ratios of  $a_{ph}/Chl-a$  which are modeled by average in the six Chl- $a$  ranges. The middle panels show relationship (c) between Chl- $a$  and  $a_{dg}$  at 442 nm, (e) between Chl- $a$  and the spectral slope of  $a_{dg}$  ( $S$ ). (d) shows the relationship between Chl- $a$  and  $a_{dg}$  at 442 nm which is modified by adjusting to  $R_{rs}$  measurements and the IOP model. The lower panels (f, g, and h) are the same as the middle panels except for  $b_{bp}$  and the spectral slope of  $b_{bp}$  ( $Y$ ) instead of  $a_{dg}$  and  $S$ .

The IOP algorithms are based on the equation of remote sensing reflectance below the surface ( $r_{rs}$ ), the total absorption coefficient ( $a$ ) and the backscattering coefficient ( $b_b$ ) proposed by Gordon et al. 1988.

$$r_{rs}(\lambda) = g_1 \times u(\lambda) + g_2 \times u(\lambda)^2 \quad (7)$$

$$a(\lambda) = a_w(\lambda) + a_{ph}(\lambda) + a_d(\lambda) + a_g(\lambda) \quad (8)$$

$$b_b(\lambda) = b_{bw}(\lambda) + b_{bp}(\lambda) \quad (9)$$

with

$$u(\lambda) = b_b(\lambda) / ( b_b(\lambda) + a(\lambda) ) \quad (10)$$

where  $g_1=0.0949$  and  $g_2=0.0794$  [Lee et al., 2002].  $a_w$ ,  $a_{ph}$ ,  $a_d$ , and  $a_g$  are the absorption spectra of water, phytoplankton, detritus, and CDOM respectively.  $b_{bw}$  and  $b_{bp}$  are backscattering coefficients of water and particles. We used  $a_w$  and  $b_{bw}$  values from [Pop and Fry 1997; Kou et al., 1993; Lee et al., 2015].

Remote sensing reflectance above the surface,  $R_{rs}$  is estimated from  $r_{rs}$  using the relation from [Gordon et al., 1988, Lee et al., 2002] as:

$$R_{rs}(\lambda) = 0.529 \times r_{rs}(\lambda) / (1 - 1.7 \times r_{rs}(\lambda)). \quad (14)$$

BRDF effect is corrected by the look-up tables developed by Morel and Maritorena (2001) by considering in-situ observation time, i.e., the solar zenith angle.

Phytoplankton absorption per Chl-*a* at 442nm,  $a_{ph}(442)/\text{Chl-}a$ , is modeled by Chl-*a* from NOMAD data. Spectral shape of the  $a_{ph}(\lambda)/\text{Chl-}a$  is modeled for six  $a_{ph}(442)/\text{Chl-}a$  levels (Fig. 1(b)).

$$a_{ph}(442)/\text{Chl-}a = -1.2937 - 0.2522 \times \text{Chl-}a + 0.0443 \times \text{Chl-}a^2 \quad (15)$$

$a_{dg}$  and  $b_{bp}$  were approximated as follows.

$$a_{dg}(\lambda) = a_{dg0} \times \exp(-S \times (\lambda - 442)) \quad (12)$$

$$b_{bp}(\lambda) = b_{bp0} \times (\lambda / 442)^{-Y} \quad (13)$$

where  $a_{dg0}$  is  $a_{dg}$  at 442 nm,  $b_{bp0}$  is  $b_{bp}$  at 442 nm were derived from the NOMAD in situ measurements of  $a_{dg}$  and  $b_{bp}$  respectively (Table 1). The average of S and Y were 0.0146, and 1.18 respectively.

SGLI  $R_{rs}$  are simulated by the IOP model and spectral response of SGLI bands, 443 (with 10-nm width), 490 (10), 530 (20), 566 (20), 672 (20) nm (Uchikawa et al., 2014), and SGLI OC4 coefficients are calculated by the simulated  $R_{rs}$  and the coincident Chl-*a* in the NOMAD.

Table 1 Coefficients for the IOP model

	VN01	VN02	VN03	VN04	VN05	VN06	VN07	VN08	VN10	VN11	
$\lambda$ [nm]	380.03	412.51	443.24	489.85	529.64	566.16	672.00	672.10	763.07	866.76	
Solar irradiance [W/m <sup>2</sup> /sr/μm]	1092.14	1712.15	1898.32	1938.46	1850.96	1797.13	1502.55	1502.30	956.34	956.62	
Refractive index	1.340	1.338	1.337	1.335	1.334	1.333	1.331	1.331	1.329	1.329	
$a_w$ [m <sup>-1</sup> ]	0.00377	0.00312	0.00510	0.01338	0.04213	0.06768	0.44579	0.44611	4.69354	4.70633	
$b_{bw}$ [m <sup>-1</sup> ]	0.00472	0.00333	0.00239	0.00157	0.00112	0.00086	0.00041	0.00041	0.00014	0.00014	
$a_{dg0}$ [m <sup>-1</sup> ]	2.48076	1.54306	0.98477	0.49841	0.27957	0.16400	0.03501	0.03496	0.00204	0.00203	
$b_{bp0}$ [m <sup>-1</sup> ]	1.19554	1.08524	0.99705	0.88615	0.80830	0.74718	0.61056	0.61045	0.45232	0.45210	
$a_{ph0}$ [m <sup>-1</sup> ]	0.65516	0.84190	0.98772	0.61764	0.30408	0.14737	0.58058	0.58094	0.00000	0.00000	
$a_{ph}/\text{Chl-}a$ dange	0.01-0.025	0.67211	0.85996	0.98719	0.63009	0.29570	0.13827	0.45687	0.45728	0.00000	0.00000
	0.025-0.04	0.66577	0.85133	0.98769	0.63673	0.27034	0.12463	0.37140	0.37173	0.00000	0.00000
	0.04-0.06	0.63217	0.81385	0.99088	0.63282	0.21233	0.08637	0.24582	0.24605	0.00000	0.00000
	0.06-0.09	0.62891	0.80775	0.99184	0.60805	0.17812	0.07189	0.19823	0.19841	0.00000	0.00000
	0.09-0.13	0.61897	0.79837	0.99323	0.61663	0.18184	0.08614	0.22302	0.22325	0.00000	0.00000
	0.13-0.5	0.64207	0.82440	0.99044	0.62424	0.22110	0.09680	0.28495	0.28521	0.00000	0.00000

## 4. Results

### 4.1. Chl-*a* estimation

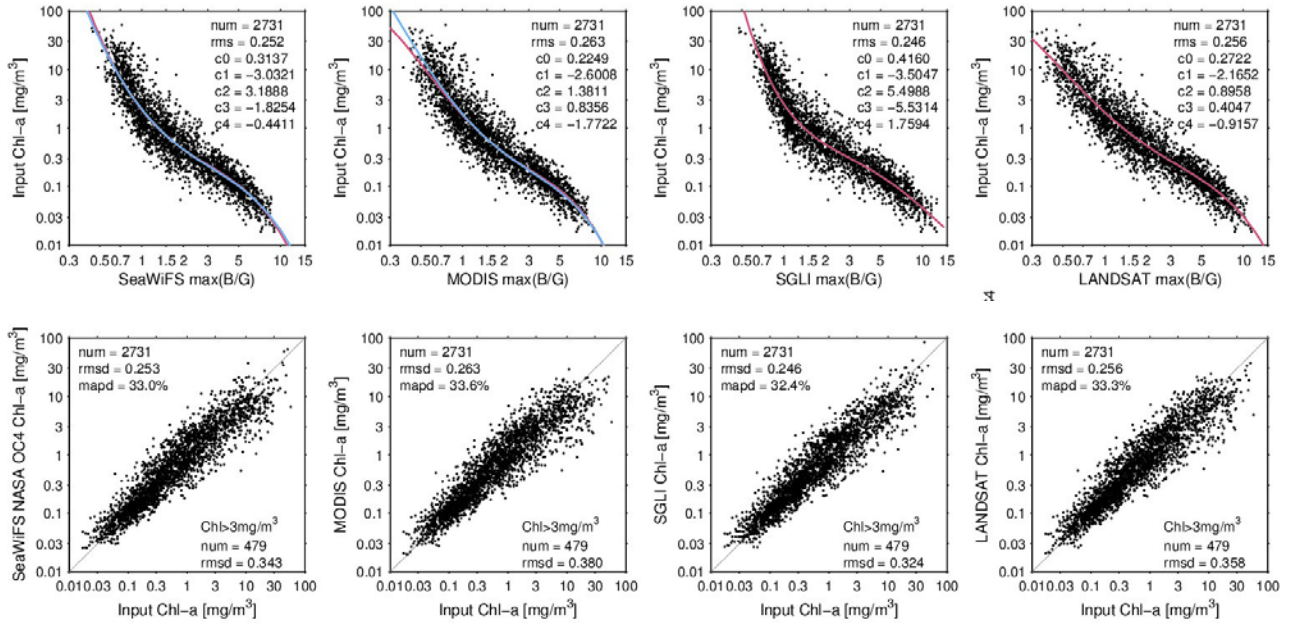


Figure 2 OC4 coefficients of SeaWiFS (NASA standard coefficients), MODIS, SGLI, and LANDSAT (upper panels), and Chl-*a* derived by the coefficients (lower panels).

The followings are SGLI Chl-*a* estimation.

$$\text{Chl-}a = \text{chl}_{e1} w_{ci} + \text{chl}_{e2} (1 - w_{ci})$$

$$w_{ci} = \frac{(-0.0002) - c_i}{(-0.0002) - (-0.0006)}$$

$$w_{ci} = 1 \quad (w_{ci} > 1)$$

$$w_{ci} = 0 \quad (w_{ci} < 0)$$

$$c_i = Rrs(\lambda_6) - (Rrs(\lambda_3) \times (\lambda_7 - \lambda_6) + Rrs(\lambda_7) \times (\lambda_6 - \lambda_3)) / (\lambda_7 - \lambda_3);$$

$$\log_{10}(\text{chl}_{e1}) = -0.38006 + 238.05110 \times c_i$$

$$\log_{10}(\text{chl}_{e2}) = 0.40451 - 3.42411 \times x + 5.29717 \times x^2 - 5.33247 \times x^3 + 1.68959 \times x^4$$

$$x = \log_{10}(\max(Rrs(\lambda_3, \lambda_4, \lambda_5)) / Rrs(\lambda_6))$$

Derived OC4 coefficients and regression errors are listed in Tables 2 and 3 respectively. RMSD of SGLI Chl-*a* seems smaller than other sensors (Fig. 2) in Chl-*a* > 3mg/m<sup>3</sup>. That is because SGLI has longer wavelengths at 530 nm and 565 nm which has higher sensitivity for the high Chl-*a* ranges.

Table 2 OCx coefficients derived in this study

	c <sub>0</sub>	c <sub>1</sub>	c <sub>2</sub>	c <sub>3</sub>	c <sub>4</sub>
<b>SGLI</b>	<b>0.40451</b>	<b>-3.42411</b>	<b>5.29717</b>	<b>-5.33247</b>	<b>1.68959</b>
SeaWiFS	0.31544	-2.95833	2.65312	-0.76475	-1.07165
MODIS	0.2249	-2.6008	1.3811	0.8356	-1.7722
LANDSAT	0.2722	-2.1652	0.8958	0.4047	-0.9157

Table 3 Chl-*a* regression error by the OCx coefficients

		0.02 < Chla < 60 N=2731	Chla < 0.1 N=356	0.1 < Chla < 3 N=1896	Chla > 3 N=479
SGLI	RMSD (log <sub>10</sub> (mg/m <sup>3</sup> ))	0.2456	0.1995	0.2301	0.3236
	MAPD(%)	32.36	27.34	31.86	39.49
SeaWiFS	RMSD (log <sub>10</sub> (mg/m <sup>3</sup> ))	0.2526	0.1922	0.2351	0.3430
	MAPD(%)	32.97	26.36	32.41	42.71
MODIS	RMSD (log <sub>10</sub> (mg/m <sup>3</sup> ))	0.2628	0.1977	0.2358	0.3803
	MAPD(%)	33.59	27.59	31.96	44.62
LANDSAT	RMSD (log <sub>10</sub> (mg/m <sup>3</sup> ))	0.2557	0.1989	0.2331	0.3580
	MAPD(%)	33.25	26.10	31.95	44.06

## 4.2. Error evaluation

Errors were estimated by add the regression errors (RMS of the regressions in Fig. 1) to the IOP model. (a)+(b)+(c) in Table 4 show errors from the in-water model, and (d), errors from the sensor noise plus the atmospheric correction.

Table 4 SGLI Chl-*a* estimation error estimated by the IOP model

	Input random error (RMS)	log <sub>10</sub>	0.02 < Chla < 60	Chla < 0.1	0.1 < Chla < 3	Chla > 3
		%	N=5000	N=991	N=2151	N=1858
(a-1) a <sub>ph</sub> /Chla [m <sup>-1</sup> /(mg/m <sup>3</sup> )]	0.1723 (log <sub>10</sub> )	RMS	0.1166	0.0679	0.1197	0.1325
		MAPD	16.30	10.48	19.76	16.79
(b-1) a <sub>dg</sub> [m <sup>-1</sup> ]	0.3978 (log <sub>10</sub> )	RMS	0.1794	0.2298	0.2082	0.0910
		MAPD	21.31	26.29	29.86	10.08
(b-2) S [μm <sup>-1</sup> ]	0.0029	RMS	0.0324	0.0064	0.0255	0.0452
		MAPD	2.69	1.04	2.55	4.68
(c-1) b <sub>bp</sub> [m <sup>-1</sup> ]	0.1924 (log <sub>10</sub> )	RMS	0.0545	0.0631	0.0618	0.0379
		MAPD	7.90	9.28	9.99	5.75
(c-2) Y	0.4724	RMS	0.0496	0.0250	0.0454	0.0626
		MAPD	6.42	3.86	6.18	9.34
(a)+(b)+(c)	-	RMS	0.2279	0.2456	0.2531	0.1830
		MAPD	30.30	30.61	36.74	25.14
(d-1) NEdL (VIS) [W/m <sup>2</sup> /sr/μm]	0.03-0.06	RMS	0.0397	0.0190	0.0298	0.0551
		MAPD	4.19	2.80	3.30	7.41
(d-2) NEdL (NIR) [W/m <sup>2</sup> /sr/μm]	0.02	RMS	0.0446	0.0360	0.0320	0.0589
		MAPD	5.39	5.15	4.08	7.86
(d-3) NEdL +Atmcorr* (Rrs 443nm)	10%	RMS	0.3761	0.6600	0.1396	0.3550
		MAPD	31.35	49.77	21.98	47.10

\* N of (d-3) are 4734, 972, 2140, 1622.

MAPD means median absolute percent difference.

## 5. Summary and discussion

SGLI Chl-*a* algorithm was developed based on O'Reilly et al., 2000 and Hu et al., 2012 and optimized by an IOP model adjusted to the NOMAD data. The OC4 and OCI coefficients are follows.

**OC4: 0.40451, -3.42411, 5.29717, -5.33247, 1.68959**

**OCI: -0.38006, 238.05110**

The coefficients will be evaluated when the standard in-situ data set is revised in the future. The IOP model constructed in this study can be used to estimate IOPs,  $a_{ph}$ ,  $a_{dg}$ , and  $b_{bp}$ .

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## Acknowledgement

NOMAD data were constructed and provided by participants in the NASA SIMBIOS Program (NRA-96-MTPE-04 and NRA-99-OES-09) and by voluntary contributors.